

Antarctic Meteorite NEWSLETTER

A periodical issued by the Antarctic Meteorite Working Group to inform scientists of the basic characteristics of specimens recovered in the Antarctic.

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!!!!!!! SAMPLE REQUEST DEADLINE: MARCH 27, 1987 (SEE PAGE 2) !!!!!!!!

	PAGE
SAMPLE-REQUEST GUIDELINES	2
METEORITE NOTES	3
NEW METEORITES	
Table 1: Alphabetical List of New 1983-1985 Specimens	6
Table 2: New 1983-1985 Specimens Listed By Type	9
Table 3: Tentative Pairings for New Specimens	11
Petrographic Descriptions	12

ATTACHMENT:

"Antarctic Meteorites: A Progress Report," by M. E. Lipschutz and W. A. Cassidy. Reprinted from <u>EOS</u> <u>67</u>, n47, p.1339-1341.

METEORITE NOTES

In case you haven't already heard, there's a new Associate Curator for Antarctic Meteorites. Last fall Marilyn Lindstrom replaced Jim Gooding in that post. Jim is now working on advanced planning for a Mars Sample Return Mission. Both Jim and Marilyn will continue their research on planetary materials.

NEWS FROM THE 1986-87 ANSMET TEAM

The six member ANSMET team returned this year to Lewis Cliff in the Beardmore Glacier area of Antarctica. Good weather aided their search and approximately 570 specimens were collected. Most of these are small ordinary chondrites, a few are carbonaceous chondrites and achondrites. The majority of the specimens were found in a 200 x 15 meter area that contained many snow drifts. The team affectionately nicknamed this area Meteorite Moraine as they worked on their hands and knees picking up the meteorites.

A1-26 and TL SURVEYS FOR METEORITE TERRESTRIAL AGES

The Meteorite Working Group has recently approved two types of surveys of Antarctic meteorites which will identify meteorites with particularly short or long terrestrial ages or unusual thermal or radiation histories. The first is gamma-ray counting for Al-26. This technique is the standard means of determining terrestrial ages and is applied by several meteorite PIs (Bhandari, Herpers, Heydegger) to analyze small numbers of samples. John Evans (Battelle Northwest) proposes a survey of 150-200 meteorites per year. This study will contribute significantly to our database on terrestrial ages of meteorites. Evans has submitted a list of requested samples but is open to suggestions from other PIs of interesting samples for which no terrestrial age is available.

The second technique involves the natural thermoluminescence (TL) of meteorites. Derek Sears and Fouad Hasan (Univ. Arkansas) conducted a pilot study for MWG which showed that natural TL correlates well with Al-26 for most meteorites. The few samples for which the two measurements do not correlate have unusual thermal or radiation histories. The results are summarized in an article in the 17th LPSC (Hasan F.A., Haq M., and Sears D.W.G. (1986) The Natural Thermoluminescence of Meteorites-I: Twenty-three Antarctic Meteorites of Known Al-26 Content). Sears is setting up a laboratory to measure natural TL in survey mode. MWG has added TL measurement to the initial processing of meteorites beginning with the 1985 collection. We look forward to publishing the results of Evans' Al-26 and Sears' TL surveys in future issues of the Newsletter.

ARTICLE BY LIPSCHUTZ AND CASSIDY

Each copy of this issue is mailed with a companion copy (reprints courtesy of LPI) of the following article:

Lipschutz M.E. and Cassidy W.A. (1986) Antarctic Meteorites: A Progress Report, EOS 67, n47, p.1339-1341.

The article briefly summarizes for the geoscience audience various aspects of the collection, curation and study of Antarctic meteorites. For a different view of the topic aimed at the chemistry audience readers are referred to: Lipschutz M.E. (1986) The Worlds Beyond: Meteorite Studies, Analytical Chemistry, 58, 968A.

A few previous issues of the Newsletter have also been accompanied by other general interest articles on Antarctic meteorites. Authors of similar articles who would like to make general distributions of reprints are invited to contact the

editor to discuss details.

NEW METEORITES FROM 1983-1985 COLLECTIONS

Pages 6-17 contain preliminary descriptions and classifications of meteorites that were completed since publication of issue 9(3) (August, 1986). Some large (>150g) specimens (regardless of petrologic type) and all "pebble"-sized (<150g) specimens of special petrologic type (carbonaceous chondrite, unequilibrated ordinary chondrite, achondrite, etc.) are represented by separate descriptions. However, some specimens of non-special petrologic type are listed only as single line entries in Table 1. For convenience, new specimens are also recast by petrologic type in Table 2.

Macroscopic descriptions of stony meteorites were performed at NASA/JSC. These descriptions summarize hand-specimen features observed during initial examination. Classification is based on microscopic petrography and reconnaissance-level electron micro-probe analyses using polished sections prepared from a small chip of each meteorite. For each stony meteorite the sample number assigned to the preliminary examination section is included. In some cases, however, a single microscopic description was based on thin sections of several specimens believed to be members of a single fall.

Meteorite descriptions contained in this issue were contributed by the following individuals:

Roberta Score, Rene' Martinez, Cecilia Satterwhite, and Carol Schwarz Planetary Materials Laboratory NASA/Johnson Space Center Northrop Services, Inc. Houston, Texas

Dr. Brian H. Mason
Department of Mineral Sciences
U. S. National Museum of Natural History
Smithsonian Institution
Washington, DC

Table 1.

List of Newly Classified Antarctic Meteorites **

Sample Number	Weight (g)	Classification	Weathering	Fracturing	% Fa	% Fs
ALH 83012	202.7	H-5 CHONDRITE	B/C	В	18	16
EET 83300 EET 83317 EET 83323 EET 83326 EET 83342 EET 83343	115.1 119.0 140.5 112.6 148.6 125.1	H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE	C B B C B	B A B B A	18 23 23 17 23 23	16 20 20 15 20 20
ALH 84101 ALH 84102 ALH 84103 ALH 84104 ALH 84105 ALH 84108 ALH 84109 ALH 84110 ALH 84112 ALH 84113 ALH 84114 ALH 84115 ALH 84115 ALH 84120 ALH 84120 ALH 84124 ALH 84132 ALH 84134 ALH 84140 ALH 84140 ALH 84140 ALH 841418 ALH 841463	220.9 213.9 137.5 201.1 260.9 214.8 245.9 318.5 145.8 212.1 119.9 194.5 113.7 129.0 114.5 157.8 113.4 164.0 130.3 168.4 100.8 134.9	H-6 CHONDRITE H-6 CHONDRITE L-3 CHONDRITE L-3 CHONDRITE L-6 CHONDRITE H-5 CHONDRITE H-5 CHONDRITE	CBBBCBBCBBCCCC	B B B C B A B B B B B B B B B B B B B B	19 24 17 24 15 18 19 18 24 18 18 18 22 18 23 23 24 24 17	17 21 15 20 14 16 16 16 16 16 16 16 21 16 20 21 15 17 15
ALH 84164 ALH 85003 ALH 85004 ALH 85010 ALH 85011 ALH 85012 ALH 85016 ALH 85017 ALH 85018 ALH 85019 ALH 85020 ALH 85021 ALH 85021 ALH 85022 ALH 85023 ALH 85024	101.4 50.1 8.4 3.2 10.7 3.9 1412.0 2361.4 811.8 632.8 744.3 646.8 951.5 438.5 387.7	CARBONACEOUS C3CARBONACEOUS C2CARBONACEOUS C2CARBON	A/B B A/B B A/B B A/B A B B A B B B B B	A A C A B A A A A A A B B A A A B B B A A A B B B A A A A B B B A A A A B B B B A A A A B B B B B A A A B	24 1-56 0.7-28 0.6-39 0.5-18 23 24 17 28 17 17 24 18 18	20 0.5-23 3-20 .8-2.5 0.7-46 20 20 15 23 15 15 20 16 15

Table 1 (cont.).

Sample Number	Weight (g)	Classification	Weathering	Fracturing	% Fa	% Fs
ALH 85025 ALH 85026 ALH 85027 ALH 85028 ALH 85029 ALH 85031 ALH 85031 ALH 85032 ALH 85033 ALH 85034 ALH 85035 ALH 85036	713.0 817.1 370.4 325.7 388.8 619.7 200.6 424.2 249.8 343.9 420.1 231.5	H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE H-6 CHONDRITE H-6 CHONDRITE H-6 CHONDRITE H-6 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE	C A B C A/B B/C B/C C A C C	A/B A B A A A A A A A A A A A A A A A A	18 24 24 19 24 17 17 17 23 24 27	16 21 20 17 21 15 15 15 6-24 21 23 16
DOM 85501 DOM 85502 DOM 85503 DOM 85504	126.2 302.2 719.7 120.6	H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE L-4 CHONDRITE	C B A B/C	A B B A	17 24 25 24	15 21 21 18-21
GEO 85700 GEO 85701	2409.0 438.6	L-6 CHONDRITE L-6 CHONDRITE	B A	A A	24 23	20 20
GRO 85203 GRO 85204 GRO 85205 GRO 85206 GRO 85207 GRO 85208 GRO 85210 GRO 85211 GRO 85212 GRO 85213	1450.4 1754.7 999.9 2420.1 2372.1 1356.9 1126.1 246.8 355.3 342.2 4364.4	H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE H-5 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE L-6 CHONDRITE H-5 CHONDRITE H-5 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE L-6 CHONDRITE	B A/B B/C A/B A B B	B A/B A B A A A A A A/B A	18 24 25 17 24 23 25 18 19 23 23	16 21 20 15 20 20 21 16 17 16-20
LEW 85307 LEW 85314 LEW 85315 LEW 85316 LEW 85318 LEW 85321 LEW 85322 LEW 85323 LEW 85324 LEW 85325 LEW 85326 LEW 85328 LEW 85329 LEW 85332 LEW 85333 LEW 85333 LEW 85345 LEW 85353 LEW 85361	1.7 14.0 10.2 34.3 152.2 11000.0 527.0 582.0 874.4 514.1 536.9 224.7 106.8 169.6 113.7 47.9 32.2 24.5 4.2	CARBONACEOUS C2 H-5 CHONDRITE H-6 CHONDRITE H-5 CHONDRITE H-5 CHONDRITE L-6 CHONDRITE UREILITE H-6 CHONDRITE CARBONACEOUS C30 L-4 CHONDRITE EUCRITE L-6 CHONDRITE	A C C C C B B C B B C B C B C B C B C B	A A/B A B/C B B A A A A B A A B A A B A A B A A B A A B A A B A A B A A B B A A B B A A B B A A B B A A B B A A B B A A B B A A B	0.6-46 18 18 17 17 18 24 19 23 18 25 19 20 19 1-20 25 17	16 16 15 15 16 20 17 20 16 21 17 17 16 1-16 21 16 22-62

Table 1 (cont.).

Sampl Numbe		Weight (g)	Classification	Weathering	Fracturin	g % Fa	% Fs
LEW 85	365	7.5	L-4 CHONDRITE	С	A	24	20
LEW 85	387	3.8	H-5 CHONDRITE	С	Α	17	15
LEW 85	390	1.5	L-4 CHONDRITE	C	A/B	24	12-24
LEW 85	396	60.2	L-3 CHONDRITE	C	A	2-26	3-25
LEW 85	401	3.9	L-3 CHONDRITE	B/C	Α	1-28	1-20
LEW 85	3440	43.8	UREILITE	В	A/B	9	8
LEW 85	6441	10.9	HOWARDITE	В	A/B		25-48
LEW 85	471	239.2	L-6 CHONDRITE	С	Α	25	22
MIL 85	600	496.9	H-5 CHONDRITE	C	Α	. 18	15

Table 2.

Newly: Classified Specimens Listed By Type **

Achondrites

Sample Number	Weight (g)	Classification	Weathering	Fracturing	% Fa	% Fs
LEW 85353	24.5	EUCRITE	В	A/B		22-62
LEW 85441	10.9	HOWARDITE	В	A/B		25-48
LEW 85328 LEW 85440	106.8 43.8	UREILITE UREILITE	B/C B	A A/B	20 9	17 8

Carbonaceous Chondrites

Sample Number	Weight (g)	Classification	Weathering	Fracturing	% Fa	% Fs
ALH 85004	8.4	CARBONACEOUS C	2 в	С		
ALH 85010	3.2	CARBONACEOUS C	2 A/B	Α	0.7-28	3-20
ALH 85011	10.7	CARBONACEOUS C2	2 A/B	A	0.6-39	.8-2.5
ALH 85012	3.9	CARBONACEOUS C2	2 B	В	0.5-18	0.7-46
LEW 85307	1.7	CARBONACEOUS C2	2 A	. А	0.6-46	- , , , ,
ALH 85003 LEW 85332	50.1 113.7	CARBONACEOUS C3	, -	A B	1-56 1-20	0.5-23 1-16

Chondrites - Type 3

Sample Number	Weight (g)	Classification	Weathering	Fracturi	ng % Fa	% Fs
ALH 84120	129.0	L-3 CHONDRITE	A/B	Α	22	6-21
LEW 85396	60.2	L-3 CHONDRITE	C	Α	2-26	3-25
LEW 85401	3.9	L-3 CHONDRITE	B/C	Α	1-28	1-20

Table 2 (cont.).

Chondrites - Type 4

Sample Number	Weight (g)	Classification	Weathering	Fracturin	ng % Fa	% Fs
ALH 84103	137.5	H-4 CHONDRITE	В	A	17	15
ALH 85033 DOM 85504 GRO 85212 LEW 85333 LEW 85365 LEW 85390	249.8 120.6 342.2 47.9 7.5 1.5	L-4 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE L-4 CHONDRITE	A B/C B B C C	A A A/B A A	23 24 23 25 24 24	6-24 18-21 16-20 21 20 12-24

** NOTES TO TABLES 1 and 2:

"Weathering" categories:

- A: Minor rustiness; rust haloes on metal particles and rust stains along fractures are minor.
- B: Moderate rustiness; large rust haloes occur on metal particles and rust stains on internal fractures are extensive.
- C: Severe rustiness; metal particles have been mostly stained by rust throughout.

"Fracturing" categories:

- A: Minor cracks; few or no cracks are conspicuous to the naked eye and no cracks penetrate the entire specimen.
- B: Moderate cracks; several cracks extend across exterior surfaces and the specimen can be readily broken along the cracks.
- c: Severe cracks; specimen readily crumbles along cracks that are both extensive and abundant.

Table 3 summarizes possible pairings of the new specimens with each other and with previously classified specimens, based on descriptive data in this newsletter issue. Readers who desire a more comprehensive review of the meteorite pairings in the U. S. Antarctic collection should refer to the compilation provided by Dr. E. R. D. Scott, as published in issue 9(2) (June, 1986).

TABLE 3.

TENTATIVE PAIRINGS FOR NEW SPECIMENS

Carbonaceous C2:

ALH85004 with ALH83100.

ALH85010, 85011, 85012 with ALH85005.

LEW85307 with LEW85306.

Carbonaceous C30:

ALH85003 with ALH82101.

L-3 Chondrite:

LEW85396, 85401.

L-6 Chondrite:

EET83317, 83323, 83342, 83343.

H-6 Chondrite:

ALH85030, 85031, 85032.

ALH84120 Sample No.: ALH841 Weight (g): 129.0 Location: Allan Hills

Field No.: 1545

Dimensions (cm): 4x4.5x3

Meteorite Type: L3 Chondrite

Macroscopic Description: Roberta Score

Thirty percent of ALH84120 is covered with fusion crust. Areas devoid of fusion crust have a brownish-gray color. The interior of this unequilibrated chondrite is medium-gray in color and contains numerous light and dark gray rounded and irregular shaped inclusions. Oxidation is light and mainly appears as haloes around metal grains.

Thin Section (,3) Description: Brian Mason

The section shows an aggregate of chondrules (0.3-1.8 mm across) and chondrule fragments in a fine-grained matrix of olivine and pyroxene with minor amounts of nickel-iron and troilite. A variety of chondrule types is present; a barred chondrule has transparent pale brown glass between the olivine bars. Most of the pyroxene is polysynthetically twinned clinobronzite. Minor weathering is indicated by brown limonitic staining around metal grains. Microprobe analyses show most olivine of fairly uniform composition, averaging Fa22, but with a few more magnesian grains (CV FeO is 8). Pyroxene composition is more variable, Fs₆₋₂₁. The meteorite is classified as an L3 chondrite, probably L3.8-3.9.

ALH85003 Sample No.: Location: Allan Hills

Weight (g): 50.1 Field No.: 2259

Dimensions (cm): 4x3.5x2.5 Meteorite Type: C30 Chondrite

Macroscopic Description: Rene Martinez

Thick patchy fusion crust covers approximately 70% of this carbonaceous chondrite. The interior is light gray and chondrules/clasts are not distinguishable in the granular matrix. A 1 mm thick weathering rind and small patches of rust are present.

Thin Section (,4) Description: Brian Mason

The thin section shows an aggregate of small chondrules (up to 0.9 mm diameter, but most are less than 0.6 mm), chondrule fragments, and irregular aggregates set in a translucent yellowbrown matrix. Chondrules are mainly granular or porphyritic olivine. Minor amounts of nickel-iron and sulfide are present, as small grains scattered throughout the section. Microprobe analyses of olivine show a wide composition range, Fa_{1.56}, mean Fa₁₇; only a few grains of pyroxene were found, having a composition range of Fs_{0.5-23}. The meteorite is classified as a C3 chondrite of the Ornans subtype; it is so similar to ALH82101 that the possibility of pairing should be considered.

Sample No.: ALH85004 Location: Allan Hills

Weight (q): 8.4 Field No.: 2298

Dimensions (cm): ten pieces Meteorite Type: C2 Chondrite

Macroscopic Description: Rene Martinez

ALH85004 consists of ten fragments plus many fines. All

fragments are black and fine-grained.

Thin Section (,3) Description: Brian Mason

The thin section shows small chondrules, irregular aggregates, and mineral grains in a dark brown to black matrix. Practically all the chondrules, aggregates, and mineral grains consist of pale green isotropic serpentinous material; a few grains of calcite were identified. The meteorite is a C2 chondrite; the extreme degree of serpentinization is reminiscent of ALH83100.

Sample No.: ALH85010, 85011, Location: Allan Hills

85012 Field Nos.: 2294; 2653;

Weight (g): 3.2; 10.7; 3.9 2221

Dimensions (cm): 2x1x0.5; 2.5x2.5x3;

many small pieces

Meteorite Type: C2 Chondrite

Macroscopic Description: Rene Martinez

Thick fusion crust is present on some of these carbonaceous chondrite fragments and not on others. All specimens have small (<1 mm) white rounded and irregular inclusions set in a fine-grained black matrix. Each fragment of ALH85012 has small amounts of evaporite deposit present. Visible oxidation is minor.

Thin Section (ALH85005,4) Description: Brian Mason
The section consists largely of black opaque matrix, through which are scattered small mineral grains (up to 0.2 mm) and sparse chondrules and chondrule fragments. The condrules and most of the mineral grains consist of olivine, usually close to Mg₂SiO₄ in composition but with some more iron-rich. Pyroxene is less common, and is close to MgSiO₃ in composition. A few grains of calcite were noted. The meteorite is a C2 chondrite; ALH85010, 85011, and 85012 as well as ALH85007, 85008, 85009, and 85013 are very similar to ALH85005 and the possibility of pairing should be considered.

LEW85307 Location: Lewis Cliff Field No.: 3178 Sample No.:

Weight (g): 1.7

Dimensions (cm): 2x1x1

Meteorite Type: C2 Chondrite

Macroscopic Description: Rene Martinez

LEW85307 is rounded and is totally covered with smooth fusion crust. Numerous <1 mm-sized white inclusions are present in the black matrix.

Thin Section (LEW85306,3) Description: Brian Mason

The section shows numerous mineral grains and aggregates and a few small (maximum diameter 0.6 mm) chondrules in a brown to black matrix. Most of the mineral grains are olivine, usually near Mg₂SiO₄ in composition, but some are more iron-rich. Pyroxene is less abundant, and is near MgSiO₃ in composition. The meteorite is a C2 chondrite. LEW85307 as well as 85009, 85011, and 85012 are very similar to LEW85306 and probably paired with it.

Sample No.: LEW85328 Location: Lewis Cliff

Weight (q): 106.8 Field No.: 2035

Dimensions (cm): 5x4x2 Meteorite Type: Ureilite

Macroscopic Description: Cecilia Satterwhite LEW85328 is covered by black fusion crust which has well defined ablation marks and is frothy in some areas. An ablation flange is present. Two mm-sized brown crystals make up the interior of this achondrite.

Thin Section (,5) Description: Brian Mason

The section shows an aggregate of anhedral to subhedral grains (0.3-2.4 mm across) of olivine and pyroxene, in the approximate proportions of 2:1. Individual grains are rimmed by carbonaceous material. Trace amounts of finely-divided nickel-iron and troilite are present, mainly along grain boundaries. Microprobe analyses show olivine of uniform composition (Fa20) with notably high CaO content (0.3%); the pyroxene is pigeonite with composition WooFs₁₇. The meteorite is a ureilite; it is notably unshocked compared to most ureilites.

Sample No.: LEW85332 Location: Lewis Cliff

Weight (g): 113.7 Field No.: 2425

Dimensions (cm): 5.5x4x3.5 Meteorite Type: C30 Chondrite

Macroscopic Description: Rene Martinez

Dark fusion crust covers all but one exterior surface. Finegrained dark gray matrix with a few <1 mm-sized light color inclusions make up the interior of this carbonaceous chondrite.

Thin Section (,4) Description: Brian Mason

The section shows an aggregate of small chondrules (up to 1.2 mm across, but most are less than 0.5 mm), chondrule fragments, and irregular granular masses set in a translucent yellow-brown matrix. Chondrules are mainly granular or porphyritic olivine. Minor amounts of nickel-iron and sulfide are present, as small grains scattered through the matrix, sometimes concentrated around chondrule rims. Olivine shows a wide composition range, Fa₁₋₂₀, mean Fa₉; pyroxene is less abundant, and has composition range Fs₁₋₁₆. The meteorite is classified as a C3 chondrite of the Ornans subtype.

Sample No.: LEW85353 Location: Lewis Cliff

Weight (g): 24.5 Field No.: 3188

Dimensions (cm): 4x2x2 Meteorite Type: Eucrite

Macroscopic Description: Rene Martinez

This pebble retains most of its thin fusion crust. Weathering of the stone has removed some fusion crust and left a pitted surface. The interior is light gray and has a basaltic texture.

Thin Section (,3) Description: Brian Mason

The section shows a fine-grained aggregate (grains 0.1-0.4 mm) of pale brown pyroxene and colorless plagioclase, with a few opaque grains. Minor weathering is indicated by brown limonitic staining around opaque grains. Plagioclase is fairly uniform in composition, mean An₈₈. Pyroxene is largely hypersthene of mean composition Wo₂Fs₆₀, together with some augite, Wo₄₈Fs₂₂; the augite shows exsolution lamellae of hypersthene. The meteorite is an unbrecciated eucrite.

Sample No.: LEW85396 Location: Lewis Cliff

Weight (g): 60.2 Field No.: 3102

Dimensions (cm): 4.5x3.0x2.0 Meteorite Type: L3 Chondrite

Macroscopic Description: Cecilia Satterwhite

Some exterior surfaces retain the black fusion crust. Areas without fusion crust are dark brown and show some light colored chondrules/inclusions. The interior is extensively weathered, though a few small inclusions/chondrules are visible.

Thin Section (,3) Description: Brian Mason
The section shows a closely packed mass of chondrules (0.3-1.8 mm across), chondrule fragments, and irregular granular aggregates, set in a small amount of dark matrix which includes minor amounts of nickel-iron and troilite. Most chondrules consist of granular or porphyritic olivine, some with polysynthetically twinned clinopyroxene. Some weathering is indicated by the presence of a

moderate amount of brown limonite as veinlets and patches. Both olivine and pyroxene show a wide range in composition: olivine, Fa_{2.26}, mean Fa₁₃; pyroxene, Fs_{3.25}. This range of compositions indicates type 3, and the small amount of nickel-iron suggests L group; the meteorite is therefore classed as an L3 chondrite.

Sample No.: LEW85401 Location: Lewis Cliff

Weight (g): 3.9 Field No.: 3124

Dimensions (cm): 2x1.5x1

Meteorite Type: L3 Chondrite

Macroscopic Description: Roberta Score

Entire angular stone is covered with dull black and brown fusion crust. Interior is heavily oxidized but the clastic nature of LEW85401 is still visible.

Thin Section (,2) Description: Brian Mason

The section is very similar to that of LEW85396, and the same description applies to both. Olivine and pyroxene show similar ranges in composition, although the mean Fa of olivine is lower in LEW85401: olivine, Fa_{1.28}, mean Fa₇; pyroxene, Fs_{1.20}. This meteorite is therefore classified as an L3 chondrite, and the possibility of pairing with LEW85396 should be considered.

LEW85440 Location: Lewis Cliff Sample No.:

Location. Field No.: 2023 43.8 Weight (g):

Dimensions (cm): 4.5x2.5x2 Meteorite Type: Ureilite

Macroscopic Description: Roberta Score Thin black fusion crust covers sixty percent of LEW85440. Greenish-gray streaks appear on all fusion crusted surfaces. Abundant minerals all showing crystal faces and areas of heavy oxidation make up the black interior.

Thin Section (,4) Description: Brian Mason The section shows an equigranular aggregate of olivine and pyroxene, as rounded to subhedral grains 0.3-0.6 mm across. grains are rimmed with black carbonaceous material, which contains trace amounts of nickel-iron (partly weathered to brown limonite) and troilite. Microprobe analyses show olivine and pyroxene of uniform composition: olivine, Fa, (CaO 0.3%); pyroxene Wo5Fs8. The meteorite is a ureilite; it appears to be relatively unshocked.

Lewis Cliff Location: LEW85441 Sample No.:

Field No.: 3125 Weight (g): 10.9

Dimensions (cm): 3x2x1.5 Meteorite Type: Howardite

Macroscopic Description: Roberta Score Fifty percent of the exterior of this achondrite is covered with dull, frothy fusion crust. Areas devoid of fusion crust are brownish-gray and contain numerous clasts. The exposed interior is lighter and grayer in color than the exterior. Abundant clasts, polymict and monomict, are present. Concentrated areas of oxidation do exist.

Thin Section (,3) Description: Brian Mason The section has a cataclastic texture, with angular fragments of pyroxene and plagioclase up to 2 mm across in a comminuted groundmass of these minerals. The pyroxene is mainly hypersthene, but ranges widely in composition: Wo₁₋₁₀, Fs₂₅₋₄₈, with a mean of Wo_4Fs_{28} . Plagioclase is fairly uniform in composition, mean An_{93} . The presence of a significant amount of pyroxene of diogenitic composition indicates that this meteorite can be classified as a howardite.